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 Swift Observations of GRB 120218AC. A. Wolf (PSU), D. Grupe (PSU), S. D. Barthelmy (NASA/GSFC), J.A. Kennea (PSU), D.N. Burrows (PSU), M.H. Siegel (PSU), and N. Gehrels (NASA/GSFC) for the Swift Team

## 1 Introduction

At 00:49:22 UT on 2012-02-18, the Swift Burst Alert Telescope (BAT) triggered and located GRB 120218A (trigger=515277). Swift did not slew due to a Sun constraint. Swift could not slew to the BAT position until 19:14 UT on 2012 March 20 (Wolf et al., GCN Circ. 12962)

The best Swift position of this burst is the BAT position given in Barthelmy et al. (GCN Circ. 12963) with RA-2000 $=21 \mathrm{~h} 19 \mathrm{~m} 03.4 \mathrm{~s}$, and Dec- $2000=-25^{\circ} 27^{\prime} 45.4^{\prime \prime}$ with an uncertainty of $1.9^{\prime}$.

## 2 BAT Observation and Analysis

At 00:49:22 UT on 2012-02-18, the Swift Burst Alert Telescope (BAT) triggered and located GRB 120218 A (trigger $=515277$, Wolf et al., GCN Circ. 12962). Using the data set from T-239 to T+303 s , the BAT ground-calculated position is RA, Dec $=319.7641,-25.4626 \mathrm{deg}$ which is
$R A(J 2000)=21 \mathrm{~h} 19 \mathrm{~m} 03.4 \mathrm{~s}$
$\operatorname{Dec}(\mathrm{J} 2000)=-25^{\circ} 27^{\prime} 45.4^{\prime \prime}$
with an uncertainty of 1.9 arcmin , (radius, sys+stat, $90 \%$ containment). The partial coding was $4 \%$ (Barthelmy et al. GCN Circ. 12963).

The mask-weighted light curve (Figure 1) shows 3 or 4 overlapping peaks starting at T-28 s, with the brightest peak at $\mathrm{T}+3 \mathrm{~s}$, and ending at $\mathrm{T}+15 \mathrm{~s}$. $T_{90}(15-350 \mathrm{keV})$ is $27.5 \pm 1.0 \mathrm{~s}$ (estimated error including systematics).
The time-averaged spectrum from T-20.6 to $\mathrm{T}+8.9 \mathrm{~s}$ is best fit by a simple power-law model. The power law index of the time-averaged spectrum is $1.75 \pm 0.11$. The fluence in the $15-150 \mathrm{keV}$ band is $5.3 \pm 0.4 \times 10^{-6} \mathrm{erg} \mathrm{cm}^{-2}$ and the 1-sec peak flux measured from $\mathrm{T}+2.33 \mathrm{~s}$ in the $15-150 \mathrm{keV}$ band is $9.1 \pm 1.2$ photons $\mathrm{cm}^{-2} \mathrm{~s}^{-1}$. All the quoted errors are at the $90 \%$ confidence level.
The results of the batgrbproduct analysis are available at http://gcn.gsfc.nasa.gov/notices_s/515277/BA/


Figure 1: BAT Light curve of GRB 120218A.

